

# High-pressure Raman and Nd<sup>3+</sup> luminescence spectroscopy of bastnäsite-(REE)CO<sub>3</sub>F

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## ABSTRACT

Bastnäsite-(Ce), a rare earth element (REE) bearing carbonate (Ce,La,Y,Nd,Pr)CO<sub>3</sub>F, is one of the most common REE-bearing minerals and has importance from both economic and geologic perspectives due to its large REE concentration. It also provides an example of the structural interplay between carbonate groups and fluorine ions, as well as the complex bonding properties of rare earth elements. We report Raman vibrational and Nd<sup>3+</sup> luminescence (<sup>4</sup>F<sub>3/2</sub>→<sup>4</sup>I<sub>9/2</sub>, <sup>4</sup>F<sub>3/2</sub>→<sup>4</sup>I<sub>11/2</sub>, and <sup>4</sup>F<sub>5/2</sub>+<sup>2</sup>H<sub>9/2</sub>→<sup>4</sup>I<sub>9/2</sub>) spectra of natural bastnäsite-(Ce) to 50 GPa at 300 K. Two phase transitions are observed under compression. Bastnäsite-I remains the stable phase up to 25 GPa, where it undergoes a subtle phase transition to bastnäsite-II. This is likely produced by a change in symmetry of the carbonate ion. Bastnäsite-II transforms to bastnäsite-III at ~38 GPa, as demonstrated by changes in the luminescence spectra. This second transition is particularly evident within the <sup>4</sup>F<sub>3/2</sub>→<sup>4</sup>I<sub>9/2</sub> luminescent transitions, and it appears that a new rare earth element site is generated at this phase change. This transition is also accompanied by modest changes in both the Raman spectra and two sets of luminescent transitions. Despite these transformations, the carbonate unit remains a stable, threefold-coordinated unit throughout this pressure range, with a possible increase in its distortion. Correspondingly, the rare-earth element site(s) appears to persist in quasi-ninefold coordination as well, implying that the general bonding configuration in bastnäsite is at least metastable over a ~30% compression range. All pressure-induced transitions are reversible, with some hysteresis, reverting to its ambient pressure phase on decompression.

**Keywords:** Bastnäsite, deep Earth, high pressure, Raman spectroscopy, rare earth elements, neodymium luminescence spectroscopy, carbon; Volatile Elements in Differentiated Planetary Interiors

## INTRODUCTION

Bastnäsite-(Ce) is a rare earth element (REE) bearing fluorocarbonate (Ce,La,Y,Nd,Pr)CO<sub>3</sub>(F,OH) that is typically found associated with carbonatite deposits and is economically important due to its high content of REEs. Indeed, bastnäsite is one of the two major mineral ores for REEs, with its chief competition being monazite. Geochemically, REEs are typically incompatible in silicate minerals; their concentration in carbonatite melts (often with bastnäsite as their crystallization product) is generally more than three orders of magnitude higher than the bulk Earth content (Jones et al. 2013), and greater than four orders of magnitude higher than within chondrites (Holtstam and Andersson 2007). Indeed, carbonate-bearing liquids and solids likely play a primary role in transporting and retaining rare earth elements within the Earth's mantle (e.g., Collerson et al. 2010). In addition, light REE have long been used to determine the ages of continental rocks, particularly via the samarium-neodymium dating method (McCulloch and Wasserburg 1978), and the evolution of the depleted mantle has been probed using the distinct <sup>143</sup>Nd/<sup>144</sup>Nd variation in the mantle compared to chondrites and crustal rocks (DePaolo and Wasserburg 1976). Bastnäsite has significance beyond its REE content: as carbonates are thought to be the dominant species containing carbon within oxidized regions of the mantle (e.g., Brenker et al. 2007), the properties

of bastnäsite can provide insight into the bonding of carbon at depth. Specifically, bastnäsite has importance for understanding the potential behavior of REE at depth within carbon (and halogen) enriched zones of the mantle, and within carbonatite melts (which are frequently halogen-enriched; e.g., Jago and Gittins 1991; Williams and Knittle 2003). Additionally, the optical and magnetic properties of rare-earth carbonate fluorides have attracted interest within the materials science community (e.g., Grice et al. 2007).

Bastnäsite virtually never occurs as an end-member in terms of its REE content, as it commonly contains a mixture of REEs (e.g., Ce, Nd, Eu, Y, etc.). Visible and infrared reflectance spectroscopy has been used to probe the absorption bands of bastnäsite, which can constrain the REE content (Turner et al. 2015); however, assignments to the absorptions and energy levels of different REEs have generally not been made to the observed absorption peaks. In addition, the rare earths in bastnäsite luminesce strongly; as discussed below, it appears that this fluorescence has at times been associated with hydroxyl stretching vibrations in Raman spectroscopic experiments (Frost and Dickfos 2007). In this study, we are able to assign the luminescence peaks to transitions of the neodymium ion, and probe changes to the REE site with pressure by using this Nd<sup>3+</sup> luminescence.

Luminescence has been employed extensively as a probe for pressure-induced changes in crystal structure or in electronic configuration (e.g., Dolan et al. 1986; Freire et al. 1994; Bray 2001; O'Bannon and Williams 2016). Neodymium luminescence has been investigated for use as a pressure calibrant due

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† Special collection papers can be found online at <http://www.minsocam.org/MSA/AmMin/special-collections.html>.

to its high fluorescence intensity and insensitivity to changes in temperature (e.g., Hua et al. 1996), with YAlO<sub>3</sub> being typically viewed as the most promising Nd-bearing sensor due to its lack of phase change or amorphization up to 80 GPa (Hua and Vohra 1997; Hernández-Rodríguez et al. 2018a, 2018b). Other high-pressure studies using Nd<sup>3+</sup> luminescence have investigated the pressure-shifts of different energy levels and their associated crystal field parameters (e.g., Muñoz-Santiuste et al. 2017; Hernández-Rodríguez et al. 2018a). Thus, Nd<sup>3+</sup> luminescence has been investigated under pressure in a range of oxide crystals to derive insights into the electronic properties of crystals used for solid state lasers under compression, and at ambient pressure within minerals (e.g., Lenz et al. 2013). However, to our knowledge, high-pressure Nd<sup>3+</sup> luminescence spectroscopy has not been previously utilized to probe the electronic and structural changes within a geologically relevant mineral. Here, we employ both Raman and Nd<sup>3+</sup> luminescence spectroscopy to constrain changes in the bonding environments of both the carbonate and neodymium ions in bastnäsite; this is the first high-pressure study of bastnäsite using these complementary techniques.

### EXPERIMENTAL METHODS

Natural bastnäsite (Ce<sub>0.49</sub>La<sub>0.25</sub>Nd<sub>0.21</sub>Pr<sub>0.04</sub>)CO<sub>3</sub>[F<sub>0.95</sub>(OH)<sub>0.05</sub>] from Northern Pakistan was used for these experiments. The chemical composition of the sample was confirmed with four grains of bastnäsite using a ThermoScientific Apreo scanning electron microscope equipped with an Oxford Instruments Ultim Max 100 mm energy-dispersive X-ray spectrometer (0.8 nA and 15 kV beam); SEM results are shown in Supplemental<sup>1</sup> Table S1 and Figure S1. Raman spectroscopy and single-crystal X-ray diffraction of this sample were in excellent agreement with previous determinations (Ni et al. 1993; Frost and Dickfos 2007).

The ambient pressure crystal structure was measured at beamline 12.2.1 at the Advanced Light Source. Diffraction images were collected at 100 K with a

Bruker D8 diffractometer equipped with a Photon II CPAD detector, and X-rays monochromated by silicon (111) to 17 keV (0.7288 Å). The single crystal was a clear block that was free of visible cracks and/or imperfections, with an approximate size of 90 × 80 × 50 μm. The sample was mounted in Paratone-N oil on a MiTeGen Micromount. Diffraction images were collected using the APEX3 program (Bruker 2016a). Images were integrated and cell refinements were computed using SAINT V8.34A (Bruker 2013). The sample symmetry of the ambient structure was determined using XPREP, and crystal absorption was corrected using the algorithms of SADABS-2016/2 (Bruker 2016b). The crystal structure from Ni et al. (1993) was used as an initial fit, and refined using full-matrix least-squares on F<sup>2</sup> via SHELXL-2018/1 (Sheldrick 2015), using the software package SHELXL Rev 806 (Hübschle et al. 2011).

Raman and luminescence spectra were collected using a Horiba LabRam HR Evolution spectrometer equipped with either a 633 or 532 nm excitation laser, depending on the experiment being conducted. The spectrometer focal length was 800 mm, and a grating of 1800 lines/mm was used with a spectral resolution of ~1 cm<sup>-1</sup>. Spectra were collected up to a pressure of ~50 GPa at ~300 K. Peak positions were calculated from fits that used a combination of Gaussian and Lorentzian peak shapes with Horiba LabSpec6 software.

High pressures were generated using a Princeton type symmetric diamond-anvil cell equipped with type IIa diamonds with 250 μm culets. Rhenium was used as the gasket material. Gaskets were preindented to ~30 μm thickness, and the sample compartments were ~125 μm in diameter. Neon was used as a pressure medium, as it is closer to hydrostatic than most other pressure media at high pressures (Klotz et al. 2009); however, supplementary data were collected using a 4:1 methanol:ethanol pressure medium, as reported in Supplemental<sup>1</sup> Figure S2. These are completely consistent with the data reported here in a neon medium. Ruby fluorescence from multiple ruby grains within the sample compartment was used to determine both pressures and pressure variations within the chamber (Dewaele et al. 2008); our results are compatible with those of Klotz et al. (2009) in there being an order of a few tenths of GPa of pressure variations at the highest pressures of these measurements.

### RESULTS

The ambient-pressure, low-temperature crystal structure (Fig. 1) of our sample agrees with the previously reported struc-

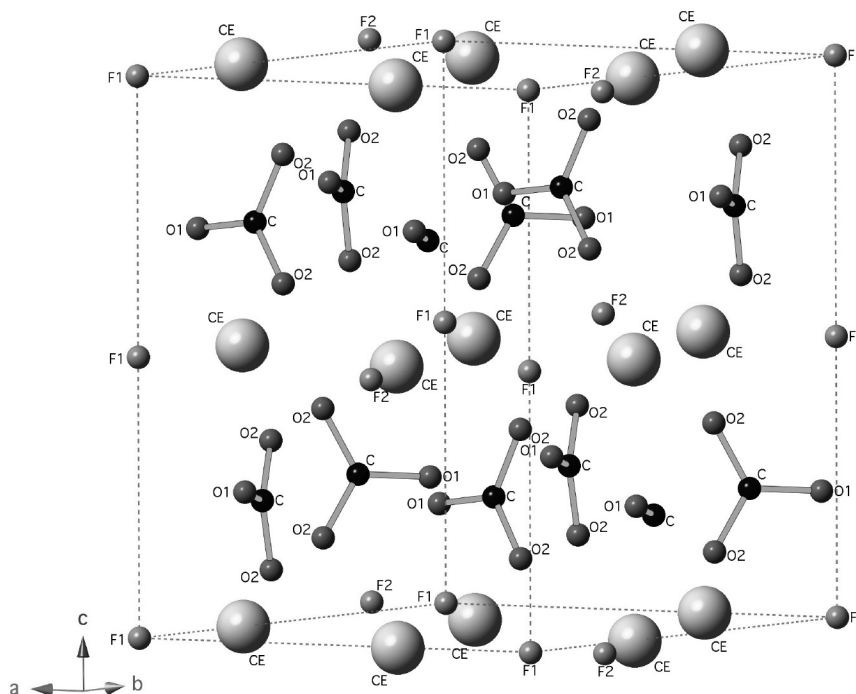


FIGURE 1. The ambient crystal structure of bastnäsite with the unit cell delineated by black lines; see labels for atom assignment.

ture by Ni et al. (1993), rather than with the structure inferred by Donnay and Donnay (1953). Bastnäsite is in the space group  $P6_2c$ , with  $a = 7.0838(2)$  Å,  $c = 9.7365(3)$  Å, and six formula units in the unit cell. Its structure has alternating layers comprised of (CO<sub>3</sub>)<sup>2-</sup> units and REE ions juxtaposed with F ions in the  $c$ -axis direction (Fig. 1). The REE atoms are in a ninefold-coordinated site (with a volume of 31.338 Å<sup>3</sup>), with a point symmetry of  $C_2$ . The atoms comprising the REE polyhedra are three fluorine atoms and six oxygen atoms. There is one carbonate group within the unit cell; the carbon and one of its coordinating O atoms (O1) lie on a mirror plane, and the other O atoms (O2) in the carbonate group are related by symmetry. The C-O1 distance is 1.2901(45) Å and C-O2( $\times 2$ ) are 1.2845(27) Å, and the carbon atom lies (within error) in a plane with the three oxygen atoms [0.0024(27)] Å out of the plane generated by the three oxygen atoms. There are two fluorine atom sites within the unit cell; F1 lies (roughly) in a plane with the REE atoms that generate a plane at  $x \sim 0.66$ . F2 lies out of this plane, causing a distortion of the REE polyhedra.

At ambient pressure and temperature, 13 lattice modes and seven modes associated with the carbonate ion are resolved in the Raman spectrum (Fig. 2a). There are a large number of

modes in the lattice region of the ambient spectrum, reflecting the relatively low symmetry of the polyhedra of the REE ions, with bonding to both oxygen and fluorine. The modes associated with the carbonate ions are easily assigned: the symmetric stretch ( $\nu_1$ ) is at 1096 cm<sup>-1</sup>, one mode associated with the out-of-plane bend ( $\nu_2$ ) is at 870 cm<sup>-1</sup>, an asymmetric stretch ( $\nu_3$ ) is at 1432 cm<sup>-1</sup>, and the in-plane bend ( $\nu_4$ ) has components at 664, 686, 729, and 737 cm<sup>-1</sup>. The increased number of carbonate modes compared to the isolated ion is due to the distortion of the carbonate ion coupled with its low symmetry site within the crystal structure. Factor-group analysis of the optic modes of bastnäsite with the space group  $P6_2c$ , yields:

$$\Gamma_{\text{optic}} = 9A'_1(\text{R}) + 7A''_1(\text{Inactive}) + 11A'_2(\text{Inactive}) + 8A''_2(\text{IR}) + 19E'(\text{R,IR}) + 16E''(\text{R}).$$

Thus, there are 44 Raman active vibrations and 27 infrared active vibrations; accordingly, the numbers of anticipated Raman active modes and carbonate modes are substantially greater than we observe, which is not unusual given that both weak modes and accidental degeneracies are typically present in complex, low-symmetry molecular crystals. At high pressures, splitting,

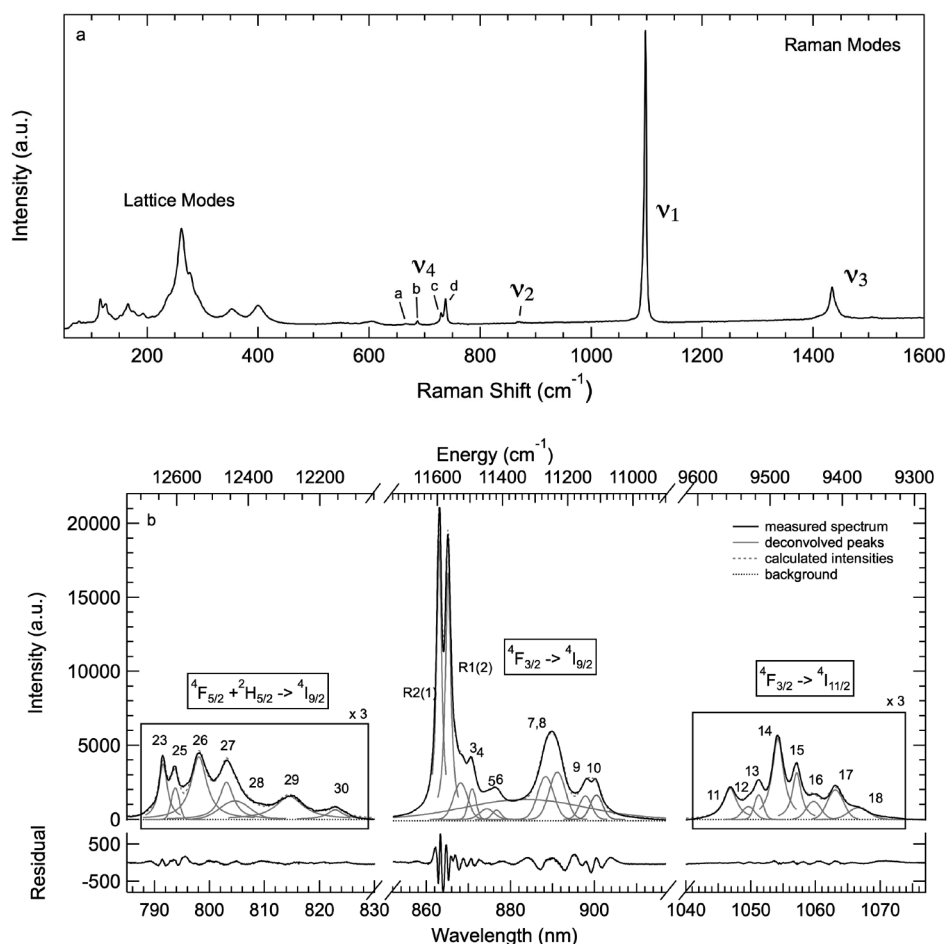


**FIGURE 2.** Ambient-pressure and -temperature (a) Raman spectrum and (b) luminescence spectrum of bastnäsite; solid black is the measured spectrum, solid gray are the deconvolved peaks, dashed gray is the calculated intensities, and dotted gray is the background. The relative intensities for the  ${}^4F_{5/2} + {}^2H_{5/2} \rightarrow {}^4I_{9/2}$  are increased by 3 $\times$ , and the  ${}^4F_{3/2} \rightarrow {}^4I_{9/2}$  section was collected for 6 $\times$  longer than the other two luminescence regions.

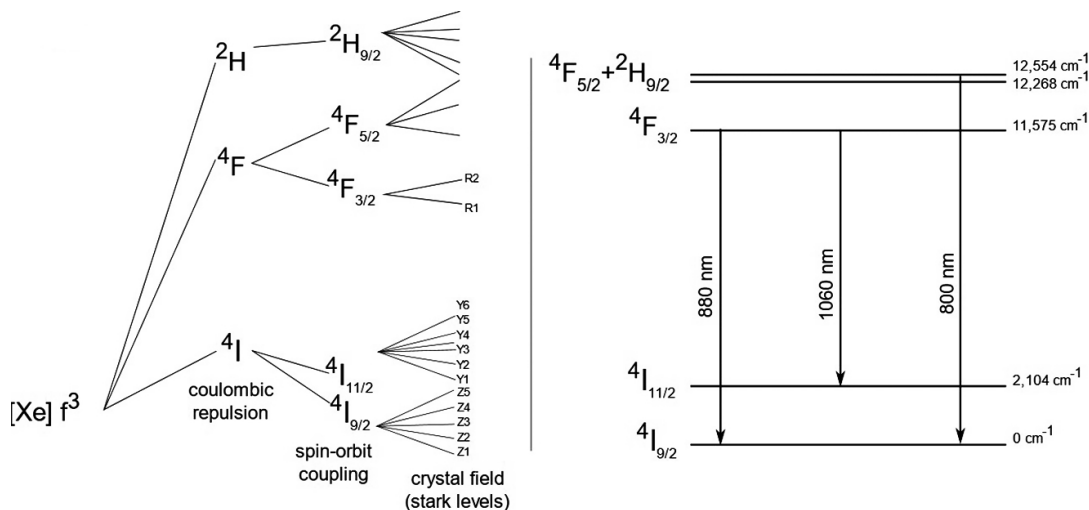


**FIGURE 3. (left)** Schematic representation of the f-electron energy levels of Nd<sup>3+</sup> in bastnäsite showing the effects of Coulombic repulsion, spin-orbit coupling, and crystal-field interactions. **(right)** Partial energy level diagram of Nd<sup>3+</sup> in bastnäsite showing the major spin-orbit coupling induced energy levels.

and mode discontinuities in the Raman data provide prima facie evidence for phase transitions that occur under compression.

Bastnäsite does not naturally occur as the pure end-member with a given rare earth element, so multi-element occupancy of the rare earth element site is expected. In our bastnäsite sample, neodymium substitutes at the 24% level into the REE site. As such, we probe the Nd<sup>3+</sup> luminescence within the structural environment of the REE site (Fig. 2b), and monitor systematic changes in site distortion via the luminescence spectroscopy. Within the bastnäsite crystal structure, neodymium's electronic energy levels are split by three principal effects: (1) by electrostatic interactions, (2) spin-orbit coupling, and (3) the local electric field that generates 2 to 6 Stark components (depending on the energy level), as imposed by the crystal field surrounding the ion (Fig. 3). The observed Nd<sup>3+</sup> luminescence is entirely consistent with that of a Nd<sup>3+</sup> ion in a site with low symmetry (e.g., Table 4.3 of Kaminskii 1981). We observe luminescence associated with energy transitions between these electronic states:  $^4F_{5/2} + ^2H_{9/2} \rightarrow ^4I_{9/2}$  at ~790–822 nm,  $^4F_{3/2} \rightarrow ^4I_{9/2}$  at ~862–900 nm, and  $^4F_{3/2} \rightarrow ^4I_{11/2}$  at ~1046–1066 nm (Fig. 3). We are able to deconvolve the peaks and assign the energy levels of each Stark-split level ( $^4I_{9/2}$ ,  $^4I_{11/2}$ ,  $F_{3/2}$ ,  $^4F_{5/2}$ ,  $^2H_{9/2}$ ) of the spin-orbit coupled energy levels (Fig. 2b).

## HIGH-PRESSURE RAMAN RESULTS AND DISCUSSION

### Bastnäsite-I: Raman spectroscopy under pressure

We resolve mode shifts for both lattice and carbonate modes of the ambient pressure phase of bastnäsite up to 25 GPa (Fig. 4). Mode assignments and their pressure shifts are reported with associated mode Grüneisen parameters in Table 1, calculated using the bulk modulus of 105 GPa from Rowland (2017). Below 25 GPa, the lattice modes all increase in frequency with pressure, with no major changes in their number or peak shape morphology (Figs. 4a and 4d). Some apparent changes in intensity are observed in this pressure range, but these changes are likely due to a combination of

pressure-induced removal of accidental degeneracies and possible effects of changes in crystal orientation between different spectra.

The Raman active modes of the carbonate ion shift monotonically up to 25 GPa, indicating that no phase change occurs up to this pressure. The symmetric stretching mode ( $\nu_1$ ) shifts approximately linearly up to 25 GPa at a rate of 2.79 cm<sup>-1</sup>/GPa. Interestingly, at ambient pressure, this peak is more asymmetric than at 23 GPa. The carbonate out-of-plane bend ( $\nu_2$ ) shifts negatively at -0.29(8) cm<sup>-1</sup>/GPa; such a modest negative shift is typical for carbonate minerals (e.g., aragonite and dolomite: Kraft et al. 1991; Vennari and Williams 2018). This mode grows in intensity on compression compared to the in-plane bend, and at 25 GPa, the mode has twice its relative amplitude as at lower pressures. This shift in intensity is plausibly associated with a Fermi resonance between higher-lying components of the  $\nu_4$  in plane bend and the  $\nu_2$  out-of-plane bends, as these bands approach one another under pressure (their separation changes from ~130 cm<sup>-1</sup> to near 75 cm<sup>-1</sup> at 25 GPa: Fig. 4b).

The four carbonate in-plane bends ( $\nu_{4a}$ ,  $\nu_{4b}$ ,  $\nu_{4c}$ ,  $\nu_{4d}$ ) all shift positively with pressure (Figs. 4b and 4e, Table 1), and their relative intensities also shift with pressure. The highest frequency band ( $\nu_{4d}$ ) remains the most intense up to 25 GPa. The second highest frequency mode ( $\nu_{4c}$ ), which initially is present as a shoulder, merges into the highest frequency band under pressure, and may decrease in relative amplitude: we believe this decrease may be from a combination of a compression-induced loss of Fermi resonance, coupled with a peak broadening of the most intense peak. There appears to be a general increase in relative intensity of the third-highest frequency in-plane bending mode with pressure ( $\nu_{4b}$ ), although the intensities of the two lowest frequency in-plane bending modes ( $\nu_{4a}$ ,  $\nu_{4b}$ ) relative to one another appear largely unchanged under compression. The pressure shifts of the two lower frequency bending modes ( $\nu_{4a}$ ,  $\nu_{4b}$ ) are such that they diverge from the two highest frequency in-plane bending modes ( $\nu_{4c}$ ,  $\nu_{4d}$ ), implying that the distortion of the carbonate group may increase under compression in bastnäsite.

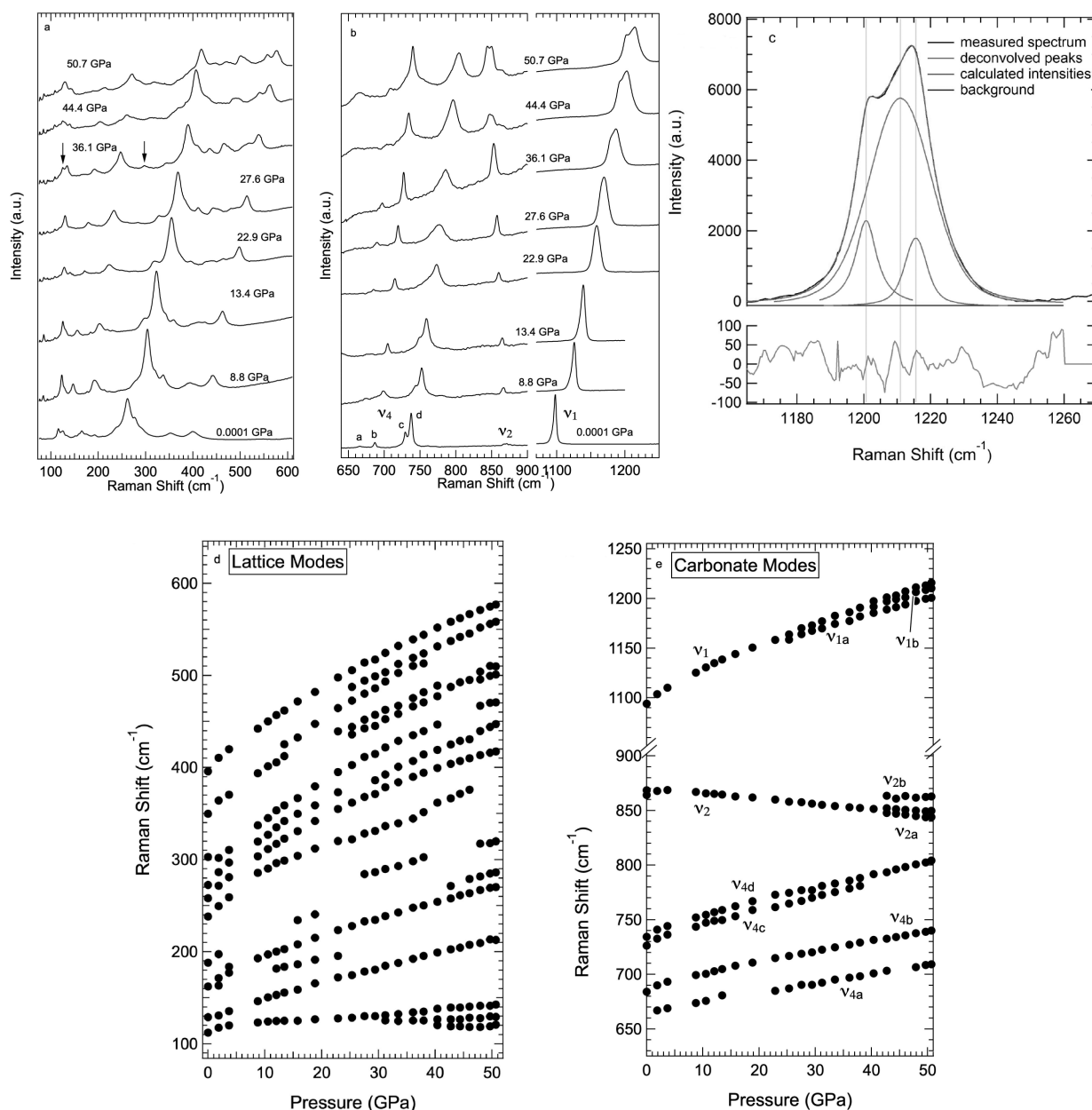


**FIGURE 4.** (a and b) Representative Raman spectra of the lattice and carbonate spectral regions under compression, arrows indicate new peaks in Raman spectrum; (c) deconvolution of the symmetric stretch of the Raman spectrum of bastnäsite at 50.7 GPa, the highest pressure probed; (d and e) peak positions of observed Raman modes as a function of pressure at room temperature under compression.

### Bastnäsite-II: Raman spectroscopy under pressure

The phase change that occurs at 25 GPa is primarily manifested by changes of the carbonate vibrations, although there are some shifts in the lattice region as well (Fig. 4). Most lattice modes appear to persist through this phase change: however, three new low-intensity modes appear at 284, 430, and 480 cm<sup>-1</sup>. Within the carbonate region, the symmetric stretch splits (Supplemental<sup>1</sup> Fig. S4), with a second peak appearing on its lower frequency side; this splitting is first resolved through deconvolution near 25 GPa (Supplemental<sup>1</sup> Fig. S4) but becomes

progressively more pronounced as pressure is increased. The two highest frequency in-plane bends (v<sub>4c</sub>, v<sub>4d</sub>), which below 25 GPa progressively merge into a single peak, become a single, asymmetric peak. The other two in-plane bend modes (v<sub>4a</sub>, v<sub>4b</sub>) and the out-of-plane bend mode each increase in intensity across the transition (Fig. 4). Taken together, these changes are consistent with a change in symmetry of the carbonate group, probably from *m* to lower symmetry.

The pressure shifts of the lattice modes in bastnäsite-II are listed in Table 2, and shown in Figures 4a and 4c. In addition

to the modes that first are resolved at 25 GPa, there are two weak modes that become unresolvable at this pressure (those with zero pressure frequencies of 140 and 185 cm<sup>-1</sup>). Relative to the carbonate region, these are relatively subtle changes in the lattice region of the Raman spectrum. Accordingly, the REE site within bastnäsite-II has likely only changed modestly compared to bastnäsite-I; certainly, no shift in the coordination of the REE site occurs. The lowest frequency lattice mode splits at 30 GPa—we believe that this may be due to pressure-induced elimination of an accidental degeneracy rather than the phase change itself, as no other lattice modes split at the same pressure. The new intermediate frequency mode (at 284 cm<sup>-1</sup> at the transition) is only readily tracked until the transition to bastnäsite-III at 38 GPa. Overall, the lattice modes undergo remarkably little alteration with pressure; the two higher frequency lattice modes that appear at 25 GPa are resolvable to the highest pressure probed. This persistence of many of the lattice modes suggests that the topological changes that occur at the two-phase transitions do not profoundly affect the lattice vibrations (which are likely more associated with the REE-F layer within the structure and vibrations of the rare earth elements against the carbonate framework). It is possible that coexistence of multiple REE sites,

with different distortions but similar coordination, occurs in the pressure range from 25 to 38 GPa.

The pressure shifts of the modes ( $\nu_1$ ,  $\nu_{1a}$ ) associated with the symmetric stretch of the carbonate ion decrease above 25 GPa, and, as pressure is increased, the two modes associated with the symmetric stretch in the high-pressure phase separate in frequency. Notably, the lower frequency in-plane bend ( $\nu_{2b}$ ) and the out-of-plane bending ( $\nu_2$ ) vibrations become sharper after the transition at 25 GPa to bastnäsite-II up until 38 GPa: this may be a consequence of the transition relieving internal stress within the crystal. Nevertheless, even though there is a change in intensity and shape of a subset of the carbonate bending modes across the transition, there is neither a change in the observed number of peaks nor a discontinuous change in their frequencies. Bastnäsite-II may have more distorted cation sites (due to the larger number of lattice peaks observed) but does not have major changes associated with its bonding, nor an apparent doubling of sites within the unit cell. The most marked effect of the transition is the splitting of the symmetric stretching vibration ( $\nu_1$ ,  $\nu_{1a}$ ) of the carbonate group, implying that the local symmetry of the carbonate group may have shifted.

### Bastnäsite-III: Raman spectroscopy under pressure

The transition from bastnäsite-II to bastnäsite-III is marked by the further splitting of the carbonate modes and modest changes in the lattice modes near 38 GPa. Again, additional modes split at this transition, and there are significant changes in band intensities: for example, the most intense lattice mode at ~390 cm<sup>-1</sup> at 38 GPa becomes markedly less intense compared to the other modes. It is unclear whether this transition, which clearly initiates at 38 GPa, is complete near this pressure, or whether it is kinetically impeded, and occurs gradually across a wide pressure range. The net decline in the intensity of several the vibrational modes of bastnäsite-II between 38 and 51 GPa, and the apparent growth of new, or weak, bands is compatible with the latter interpretation. If, however, this is a transition that is complete near 38 GPa, the new modes (at 120 and 271 cm<sup>-1</sup>) that appear at the transition indicate that either the REE site is further distorted, and/or there has been a doubling of the REE sites within the unit cell. We believe, in conjunction with the luminescence data described below, that a new REE site is produced associated with the transition to bastnäsite-III.

The changes of the carbonate vibrations at the transition to bastnäsite-III are more dramatic than those associated with the bastnäsite-I to bastnäsite-II transition. The symmetric stretch, which splits into two deconvolved modes ( $\nu_1$ ,  $\nu_{1a}$ ) above 25 GPa, requires a new component ( $\nu_{1b}$ ) to fit its band shape above 38 GPa (Fig. 4c, Table 3). The out-of-plane bend begins to become asymmetric at 36 GPa, and splits from one band to three distinct modes ( $\nu_2$ ,  $\nu_{2a}$ ,  $\nu_{2b}$ ) above 38 GPa. The highest frequency in-plane bending peak ( $\nu_{4d}$ ) sharpens across the transition to bastnäsite-III; the two components ( $\nu_{4c}$ ,  $\nu_{4d}$ ) of this peak become difficult to deconvolve above 38 GPa. The next highest frequency in-plane bending peak ( $\nu_{4b}$ ) remains sharp and distinct across the transition, but a change in its peak shape is apparent: the top of the peak remains sharp, but the bottom broadens, and there may be additional, not-readily-resolved modes on its high and low frequency sides. Accordingly, two of the three (and possibly all

**TABLE 1.** Raman modes as a function of pressure and calculated mode Grüneisen parameters of bastnäsite-I up to 25 GPa

| $\nu_0$ (cm <sup>-1</sup> ) | Assignment                                                      | $d\nu/dP$ (cm <sup>-1</sup> /GPa) | Grüneisen parameter <sup>a</sup> |
|-----------------------------|-----------------------------------------------------------------|-----------------------------------|----------------------------------|
| 112                         | Lattice mode                                                    | 0.59(9)                           | 0.53                             |
| 129                         | Lattice mode                                                    | 1.95(4)                           | 1.59                             |
| 162                         | Lattice mode                                                    | 2.53(16)                          | 1.64                             |
| 188                         | Lattice mode                                                    | 2.74(14)                          | 1.54                             |
| 238                         | Lattice mode                                                    | 3.63(28)                          | 1.60                             |
| 258                         | Lattice mode                                                    | 4.19(16)                          | 1.70                             |
| 272                         | Lattice mode                                                    | 4.36(18)                          | 1.68                             |
| 303                         | Lattice mode                                                    | 4.33(14)                          | 1.50                             |
| 350                         | Lattice mode                                                    | 3.90 (24)                         | 1.17                             |
| 396                         | Lattice mode                                                    | 4.35(14)                          | 1.15                             |
| —                           | (CO <sub>3</sub> ) <sup>2-</sup> in-plane bend ( $\nu_{4a}$ )   | 0.89(10)                          | —                                |
| 684                         | (CO <sub>3</sub> ) <sup>2-</sup> in-plane bend ( $\nu_{4b}$ )   | 1.28(6)                           | 0.20                             |
| 726                         | (CO <sub>3</sub> ) <sup>2-</sup> in-plane bend ( $\nu_{4c}$ )   | 1.51(8)                           | 0.22                             |
| 734                         | (CO <sub>3</sub> ) <sup>2-</sup> in-plane bend ( $\nu_{4d}$ )   | 1.60(6)                           | 0.23                             |
| 864                         | (CO <sub>3</sub> ) <sup>2-</sup> out-of-plane bend ( $\nu_2$ )  | —                                 | —                                |
| 868                         | (CO <sub>3</sub> ) <sup>2-</sup> out-of-plane bend ( $\nu_2$ )  | -0.29(8)                          | -0.04                            |
| 1094                        | (CO <sub>3</sub> ) <sup>2-</sup> symmetric stretch ( $\nu_1$ )  | 2.79(13)                          | 0.30                             |
| 1434                        | (CO <sub>3</sub> ) <sup>2-</sup> asymmetric stretch ( $\nu_3$ ) | —                                 | —                                |

<sup>a</sup> Bulk modulus of 105 GPa used from the La/F bastnäsite end-member (Rowland 2017).

**TABLE 2.** Raman modes as a function of pressure of bastnäsite-II from 25 to 38 GPa

| $\nu_0$ (cm <sup>-1</sup> ) <sup>a</sup> | Assignment                                                        | $d\nu/dP$ (cm <sup>-1</sup> /GPa) |
|------------------------------------------|-------------------------------------------------------------------|-----------------------------------|
| 128                                      | Lattice mode                                                      | 0.52(3)                           |
| 175                                      | Lattice mode                                                      | 1.66(4)                           |
| 228                                      | Lattice mode                                                      | 1.77(6)                           |
| 284                                      | Lattice mode                                                      | 1.76(8)                           |
| 322                                      | Lattice mode                                                      | 2.20(11)                          |
| 362                                      | Lattice mode                                                      | 2.61(9)                           |
| 403                                      | Lattice mode                                                      | 2.92(11)                          |
| 436                                      | Lattice mode                                                      | 2.82(9)                           |
| 444                                      | Lattice mode                                                      | 2.88(8)                           |
| 473                                      | Lattice mode                                                      | 3.35(15)                          |
| 488                                      | Lattice mode                                                      | 2.92(8)                           |
| 506                                      | Lattice mode                                                      | 3.06(11)                          |
| 687                                      | (CO <sub>3</sub> ) <sup>2-</sup> in-plane bend ( $\nu_{4a}$ )     | 0.87(6)                           |
| 717                                      | (CO <sub>3</sub> ) <sup>2-</sup> in-plane bend ( $\nu_{4b}$ )     | 0.98(2)                           |
| 765                                      | (CO <sub>3</sub> ) <sup>2-</sup> in-plane bend ( $\nu_{4c}$ )     | 1.39(3)                           |
| 775                                      | (CO <sub>3</sub> ) <sup>2-</sup> in-plane bend ( $\nu_{4d}$ )     | 1.09(7)                           |
| 858                                      | (CO <sub>3</sub> ) <sup>2-</sup> out-of-plane bend ( $\nu_2$ )    | -0.49(2)                          |
| 1158                                     | (CO <sub>3</sub> ) <sup>2-</sup> symmetric stretch ( $\nu_{1a}$ ) | 1.66(7)                           |
| 1164                                     | (CO <sub>3</sub> ) <sup>2-</sup> symmetric stretch ( $\nu_1$ )    | 2.09(7)                           |

<sup>a</sup>  $\nu_0$  is at 25 GPa.

three) of the different types of carbonate vibrations monitored have additional bands that appear following the transition at 38 GPa.

The Raman spectrum of bastnäsite-III reflects both distorted carbonate ions and potentially multiple REE sites that persist to the highest pressures probed (51 GPa). While Raman spectroscopy is an excellent probe for characterizing changes in local bonding environments, we supplement these results with Nd<sup>3+</sup> luminescence results to characterize the electronic transitions of Nd in the REE site to better constrain structural changes that specifically involve the REE ion.

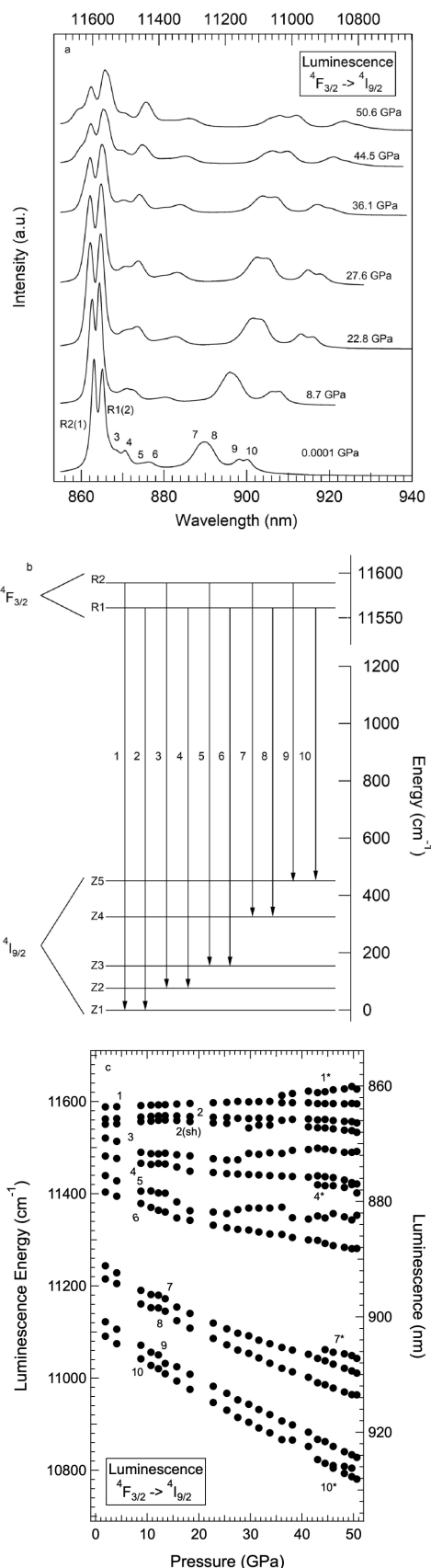
### HIGH-PRESSURE ND<sup>3+</sup> LUMINESCENCE RESULTS AND DISCUSSION

Figures 5 to 7 show both our luminescence results within different spectral regions and the ambient-pressure energy level diagrams for the associated transitions. In each instance, we use the peak nomenclature of Dieke (1969). The majority of the transitions have red shifts of their respective luminescence with pressure—such an approach of long-lived luminescing electronic states toward the ground state under compression is common, such as for Nd<sup>3+</sup> in garnets (Hua et al. 1996) and Cr<sup>3+</sup> in ruby (Mao et al. 1986), although exceptions exist (e.g., Hua and Vohra 1997; Hernández-Rodríguez et al. 2018a). The most intense and highest energy bands associated with the transition between the <sup>4</sup>F<sub>3/2</sub> state and the <sup>4</sup>I<sub>9/2</sub> ground state (which are, by convention, named R1 and R2; Dieke 1969), do not redshift or blueshift notably; while they may have a small and nonlinear pressure dependence, they shift little with pressure up to 50 GPa (Fig. 5, Table 4). Thus, the pressure-induced shifts of the <sup>4</sup>F<sub>3/2</sub> states are similar to that of the lowest energy <sup>4</sup>I<sub>9/2</sub> state.

Overall, the luminescence changes subtly at the transition between bastnäsite-I and -II compared to the changes that occur across the bastnäsite-II to -III transition; this is likely due to the subtle nature of the first transition, and a more dramatic change in the REE site in the second transition (possibly involving a doubling of the number of REE sites in the cell). We discuss each of the major transition manifolds in sequence, since each show somewhat different behavior under compression.

#### <sup>4</sup>F<sub>3/2</sub>→<sup>4</sup>I<sub>9/2</sub> luminescence under pressure

We observe all 10 peaks associated with the transitions between the Stark levels for the <sup>4</sup>F<sub>3/2</sub> and <sup>4</sup>I<sub>9/2</sub> states. The R1 and R2 bands provide a direct measure of the excited state (<sup>4</sup>F<sub>3/2</sub>) energy above the ground state. Under pressure, up until ~15 GPa, we see a decrease in their peak separation (Fig. 8). However, near 15 GPa this trend reverses, and the peaks begin to separate. Because the Stark splitting is produced by the local electric field at the Nd ion, this turnover is associated with an initial decrease in the electric field gradient, followed by an increase, or an onset of a more anisotropic local bonding environment. Such a change may be indicative of a change in local bonding, and hence a shift in compressional behavior within the REE site, but does not



► **FIGURE 5.** Nd<sup>3+</sup> luminescence of the <sup>4</sup>F<sub>3/2</sub>→<sup>4</sup>I<sub>9/2</sub> transition (a) luminescence spectra as a function of pressure. (b) Partial energy level diagram of this transition. (c) Peak positions of observed luminescence bands with pressure.

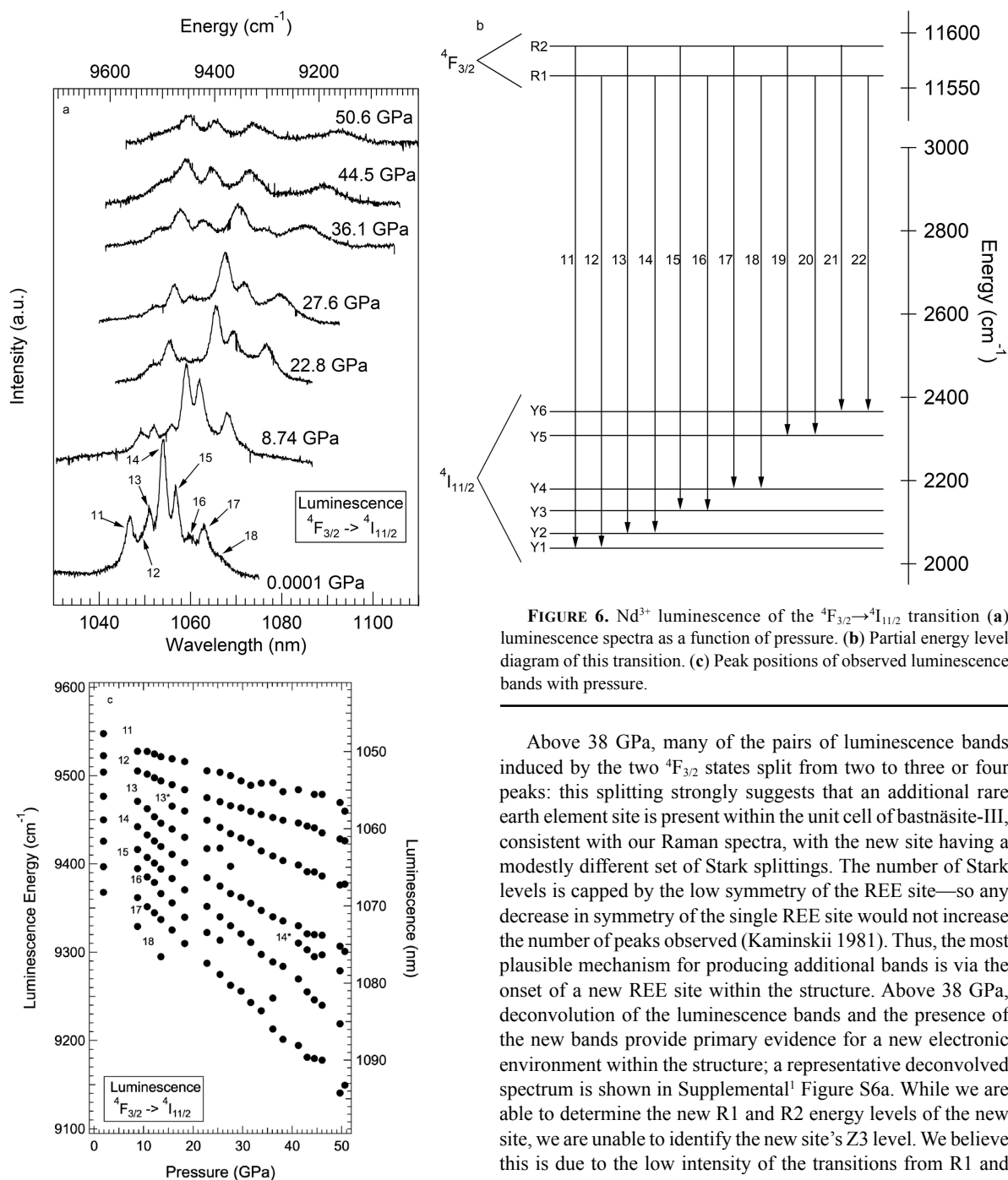


**FIGURE 6.** Nd<sup>3+</sup> luminescence of the  $^4F_{3/2} \rightarrow ^4I_{11/2}$  transition (a) luminescence spectra as a function of pressure. (b) Partial energy level diagram of this transition. (c) Peak positions of observed luminescence bands with pressure.

Above 38 GPa, many of the pairs of luminescence bands induced by the two  $^4F_{3/2}$  states split from two to three or four peaks: this splitting strongly suggests that an additional rare earth element site is present within the unit cell of bastnäsite-III, consistent with our Raman spectra, with the new site having a modestly different set of Stark splittings. The number of Stark levels is capped by the low symmetry of the REE site—so any decrease in symmetry of the single REE site would not increase the number of peaks observed (Kaminskii 1981). Thus, the most plausible mechanism for producing additional bands is via the onset of a new REE site within the structure. Above 38 GPa, deconvolution of the luminescence bands and the presence of the new bands provide primary evidence for a new electronic environment within the structure; a representative deconvolved spectrum is shown in Supplemental<sup>1</sup> Figure S6a. While we are able to determine the new R1 and R2 energy levels of the new site, we are unable to identify the new site's Z3 level. We believe this is due to the low intensity of the transitions from R1 and R2 to Z3 before the phase transition; thus, after the transition it remains difficult to determine the full suite of peak locations associated with the new site.

Interestingly, a new band begins to appear at the highest energies of this suite of transitions (associated with an additional transition from R2 → Z1) near 36 GPa—this is likely either a precursor to, or the first indication of the phase change that we observe in the Raman spectra at 38 GPa. The intensity of the peak associated with R1 → Z2(4) becomes more intense in bastnäsite-III. The change in intensity indicates that the excitation and/or emission cross sections for those transitions have changed:

necessarily indicate any change in crystal symmetry. The relative intensities of these bands remain relatively unchanged to 25 GPa.

From 25 to 38 GPa, most of the luminescence peak positions continue to decrease in energy at rates similar to those prior to the transition to bastnäsite-II; a representative spectrum at 31.2 GPa is peakfit in Supplemental<sup>1</sup> Fig. S5a. Interestingly, only one transition clearly increases in energy above 25 GPa: R2 → Z2(3). This anomalous shift indicates that the splitting of the  $^4I_{9/2}$  state is complex under compression and that the electric field that is impacting the f-orbitals is anisotropic in its shape.

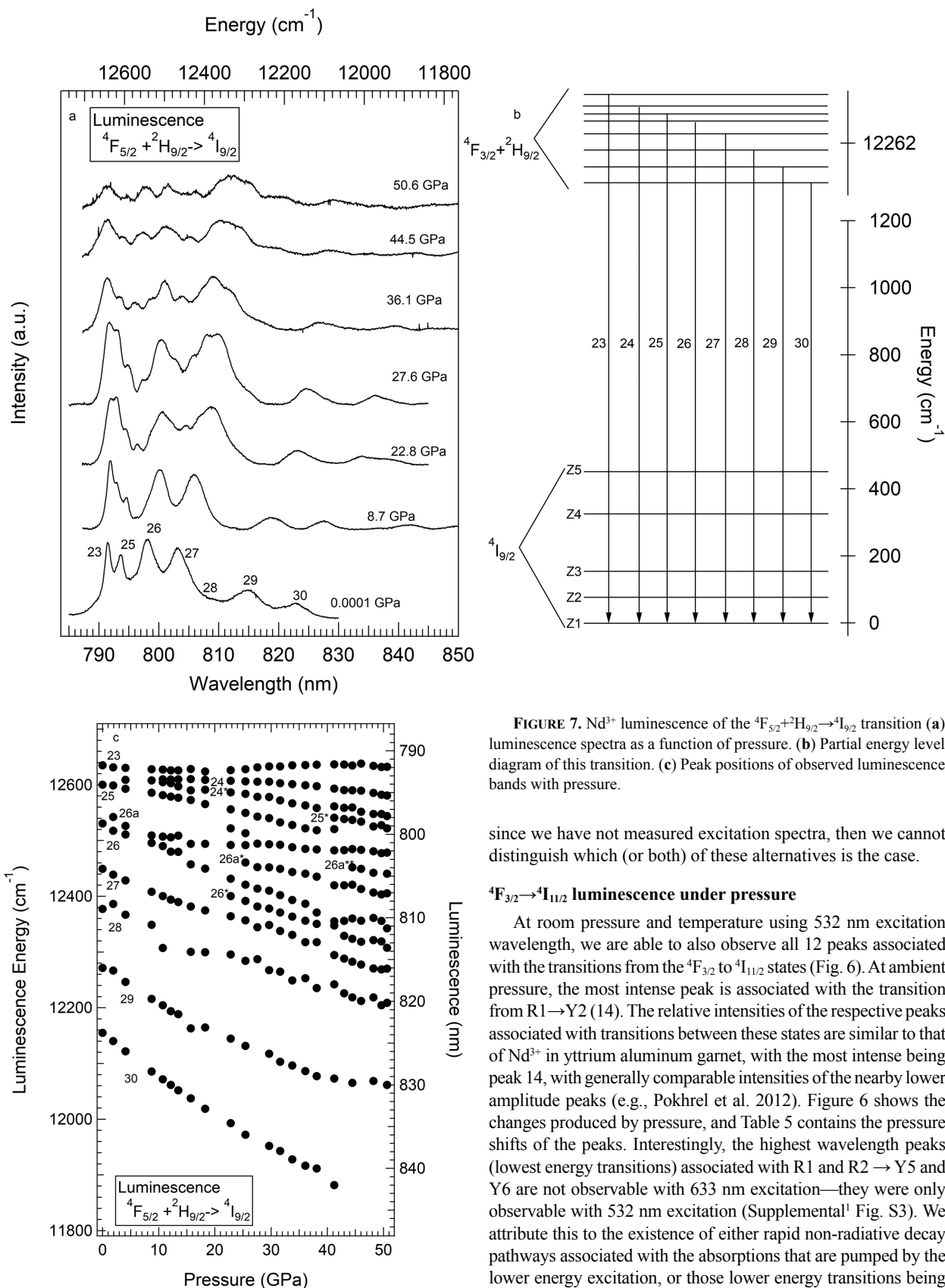


FIGURE 7. Nd<sup>3+</sup> luminescence of the  ${}^4F_{5/2} + {}^2H_{9/2} \rightarrow {}^4I_{9/2}$  transition (a) luminescence spectra as a function of pressure. (b) Partial energy level diagram of this transition. (c) Peak positions of observed luminescence bands with pressure.

since we have not measured excitation spectra, then we cannot distinguish which (or both) of these alternatives is the case.

#### ${}^4F_{3/2} \rightarrow {}^4I_{11/2}$ luminescence under pressure

At room pressure and temperature using 532 nm excitation wavelength, we are able to also observe all 12 peaks associated with the transitions from the  ${}^4F_{3/2}$  to  ${}^4I_{11/2}$  states (Fig. 6). At ambient pressure, the most intense peak is associated with the transition from R1  $\rightarrow$  Y2 (14). The relative intensities of the respective peaks associated with transitions between these states are similar to that of Nd<sup>3+</sup> in yttrium aluminum garnet, with the most intense being peak 14, with generally comparable intensities of the nearby lower amplitude peaks (e.g., Pokhrel et al. 2012). Figure 6 shows the changes produced by pressure, and Table 5 contains the pressure shifts of the peaks. Interestingly, the highest wavelength peaks (lowest energy transitions) associated with R1 and R2  $\rightarrow$  Y5 and Y6 are not observable with 633 nm excitation—they were only observable with 532 nm excitation (Supplemental<sup>1</sup> Fig. S3). We attribute this to the existence of either rapid non-radiative decay pathways associated with the absorptions that are pumped by the lower energy excitation, or those lower energy transitions being

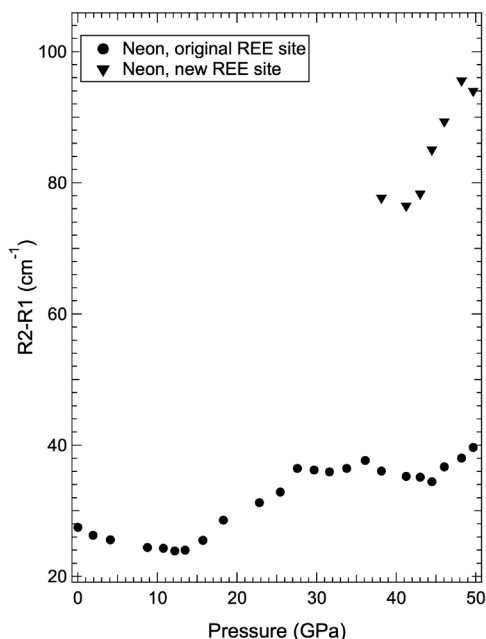


**FIGURE 8.** R-line separation of Nd<sup>3+</sup> (R<sub>1</sub> and R<sub>2</sub>) in the original REE site (closed circles), and associated with the new high-pressure site in bastnäsite-III (closed triangles), as a function of pressure.

accessible when the higher energy transitions are overpopulated.

Under compression, all modes shifted to lower energy with increased pressure, and as with the <sup>4</sup>I<sub>9/2</sub> transitions, the modes with the lowest 1 bar energy exhibited the strongest negative pressure dependence. The lowest energy mode (18) becomes unresolvable above 15 GPa. Around this pressure, an additional peak 13\* is observed at slightly lower energy than peak 13. These changes in peak position and intensity may be associated with the change in the sign of the trend of the electric field near this pressure that is observed in the R<sub>2</sub>-R<sub>1</sub> separation (Fig. 8). At 25 GPa, peak 13 becomes unresolvable, and this shift may be associated with the transition from bastnäsite-I to bastnäsite-II. Overall, across the bastnäsite-I to bastnäsite-II transition, there are no major changes in either the number of luminescence peaks or their intensities, thus confirming our inference from the Raman spectra that this transition likely involves (at most) a subtle change to the REE site. Moreover, the pressure shifts of all bands are consistent in their migration to lower energy, indicating that these transitions may be less sensitive to changes in the detailed electric field gradient.

Above 25 GPa, the peaks associated with bastnäsite-II continue to broaden and decrease in intensity. As pressure is increased, peak 14 decreases particularly in intensity relative to the other bands in this multiplet. All peaks, except peak 16 (which becomes unresolvable) persist to 38 GPa; a representative spectrum at 31.2 GPa is peakfit in Supplemental<sup>1</sup> Figure S5b.

Above 38 GPa, the relative intensities of individual peaks continue to change. As pressure is increased, the lower energy transitions become less intense relative to the higher energy transitions; specifically, peak 14, which at ambient pressure is the most intense, becomes the third most intense, following peaks 12 and 13. These intensity changes may be due to a splitting of peak 14. At 38 GPa, a lower energy peak branches off of peak 14. This splitting,

coupled with the overall drop in intensity of the multiplet of peaks, further indicates that there is likely an additional, distinct REE site in the unit cell above 38 GPa; the deconvolution of a representative spectrum at 50.6 GPa is shown in Supplemental<sup>1</sup> Fig. S6b.

#### <sup>4</sup>F<sub>5/2</sub>+<sup>2</sup>H<sub>9/2</sub>→<sup>4</sup>I<sub>9/2</sub> luminescence under pressure

At ambient pressure and temperature, we observe seven peaks associated with the transition from <sup>4</sup>F<sub>5/2</sub>+<sup>4</sup>H<sub>9/2</sub>→<sup>4</sup>I<sub>9/2</sub>. There is a well-documented overlap between the energy levels of <sup>4</sup>F<sub>5/2</sub> and <sup>4</sup>H<sub>9/2</sub> (5 in <sup>4</sup>H<sub>9/2</sub> and 3 in <sup>4</sup>F<sub>5/2</sub>), and identification of the individual Stark levels of the <sup>4</sup>F<sub>5/2</sub>+<sup>4</sup>H<sub>9/2</sub> transitions to the <sup>4</sup>I<sub>9/2</sub> state

**TABLE 3.** Raman modes as a function of pressure of bastnäsite-III from 38 to 50 GPa

| $\nu_0$ (cm <sup>-1</sup> ) <sup>a</sup> | Assignment                                                        | dv/dP (cm <sup>-1</sup> /GPa) |
|------------------------------------------|-------------------------------------------------------------------|-------------------------------|
| 119                                      | Lattice mode                                                      | 0.007(116)                    |
| 127                                      | Lattice mode                                                      | 0.32(8)                       |
| 139                                      | Lattice mode                                                      | 0.40(4)                       |
| 203                                      | Lattice mode                                                      | 1.39(8)                       |
| 258                                      | Lattice mode                                                      | 1.59(5)                       |
| 271                                      | Lattice mode                                                      | 1.84(12)                      |
| 367                                      | Lattice mode                                                      | 2.50(17)                      |
| 404                                      | Lattice mode                                                      | 1.74(4)                       |
| 425                                      | Lattice mode                                                      | 2.75(17)                      |
| 488                                      | Lattice mode                                                      | 2.07(29)                      |
| 537                                      | Lattice mode                                                      | 2.65(6)                       |
| 558                                      | Lattice mode                                                      | 2.43(5)                       |
| 703                                      | (CO <sub>3</sub> ) <sup>2-</sup> in-plane bend ( $\nu_{4a}$ )     | 0.78(4)                       |
| 733                                      | (CO <sub>3</sub> ) <sup>2-</sup> in-plane bend ( $\nu_{4b}$ )     | 0.85(4)                       |
| 794                                      | (CO <sub>3</sub> ) <sup>2-</sup> in-plane bend ( $\nu_{4d}$ )     | 1.23(5)                       |
| 848                                      | (CO <sub>3</sub> ) <sup>2-</sup> out-of-plane bend ( $\nu_{2a}$ ) | -0.68(9)                      |
| 852                                      | (CO <sub>3</sub> ) <sup>2-</sup> out-of-plane bend ( $\nu_2$ )    | -0.31(4)                      |
| 864                                      | (CO <sub>3</sub> ) <sup>2-</sup> out-of-plane bend ( $\nu_{2b}$ ) | -0.007(162)                   |
| 1186                                     | (CO <sub>3</sub> ) <sup>2-</sup> symmetric stretch ( $\nu_{1a}$ ) | 1.54(5)                       |
| 1192                                     | (CO <sub>3</sub> ) <sup>2-</sup> symmetric stretch ( $\nu_{1b}$ ) | 1.79(8)                       |
| 1197                                     | (CO <sub>3</sub> ) <sup>2-</sup> symmetric stretch ( $\nu_1$ )    | 1.81(7)                       |

Note: Asterisks indicate new peaks. <sup>a</sup>  $\nu_0$  is at 38 GPa.

**TABLE 4.** <sup>4</sup>F<sub>3/2</sub>→<sup>4</sup>I<sub>9/2</sub> luminescence as a function of pressure in bastnäsite

| Assignment | $\lambda_0$ (nm) | $\nu_0$ (cm <sup>-1</sup> ) | Bastnäsite-I to 25 GPa (cm <sup>-1</sup> /GPa) | Bastnäsite-II 25 to 40 GPa (cm <sup>-1</sup> /GPa) | Bastnäsite-II 4 to 50 GPa (cm <sup>-1</sup> /GPa) |
|------------|------------------|-----------------------------|------------------------------------------------|----------------------------------------------------|---------------------------------------------------|
| 1*         |                  |                             |                                                |                                                    | 1.05(33)                                          |
| 1          | 862.91           | 11,588.7                    | 0.42(2)                                        | -0.27(11)                                          | -0.12(7)                                          |
| 2          | 864.96           | 11,561.2                    | 0.20(9)                                        | -0.31(16)                                          | -0.83(11)                                         |
| 3          | 868.47           | 11,514.5                    | -1.84(24)                                      | 1.43(48)                                           | -1.13(13)                                         |
| 4          | 870.63           | 11,485.9                    | -1.67(10)                                      | -0.48(2)                                           | -0.87(29)                                         |
| 4*         |                  |                             |                                                |                                                    | -1.76(37)                                         |
| 5          | 874.51           | 11,435.0                    | -3.51(26)                                      | -1.14(94)                                          | -1.37(88)                                         |
| 6          | 876.80           | 11,405.1                    | -3.35(13)                                      | -1.61(15)                                          | 0.16(62)                                          |
| 7*         |                  |                             |                                                |                                                    | -2.76(36)                                         |
| 7          | 887.95           | 11,261.9                    | -5.99(23)                                      | -3.57(14)                                          | -4.28(7)                                          |
| 8          | 889.68           | 11,240.0                    | -6.40(26)                                      | -4.63(15)                                          | -4.08(24)                                         |
| 9          | 897.84           | 11,137.8                    | -6.73(19)                                      | -5.26(10)                                          | -5.62(25)                                         |
| 10         | 900.23           | 11,108.3                    | -6.95(8)                                       | -4.93(40)                                          | -4.75(134)                                        |
| 10*        |                  |                             |                                                |                                                    | -5.16(2)                                          |

Note: Asterisks indicate new peaks.

**TABLE 5.** <sup>4</sup>F<sub>3/2</sub>→<sup>4</sup>I<sub>11/2</sub> luminescence as a function of pressure in bastnäsite

| Assignment | $\lambda_0$ (nm) | $\nu_0$ (cm <sup>-1</sup> ) | Bastnäsite-I to 25 GPa (cm <sup>-1</sup> /GPa) | Bastnäsite-II 25 to 40 GPa (cm <sup>-1</sup> /GPa) | Bastnäsite-II 4 to 50 GPa (cm <sup>-1</sup> /GPa) |
|------------|------------------|-----------------------------|------------------------------------------------|----------------------------------------------------|---------------------------------------------------|
| 11         | 1046.83          | 9552.7                      | -1.91(11)                                      | -1.26(40)                                          | -2.35(51)                                         |
| 12         | 1049.86          | 9525.1                      | -2.21(4)                                       | -1.59(5)                                           | -2.19(10)                                         |
| 13*        |                  |                             | -2.47(12)                                      | -2.99(14)                                          | -2.33(22)                                         |
| 13         | 1051.0           | 9514.7                      | -4.00(27)                                      |                                                    |                                                   |
| 14         | 1054.01          | 9487.6                      | -4.42(14)                                      | -3.08(13)                                          | -2.78(33)                                         |
| 14*        |                  |                             |                                                |                                                    | -3.54(50)                                         |
| 15         | 1056.82          | 9462.3                      | -4.77(7)                                       | -4.57(27)                                          | -5.84(37)                                         |
| 16         | 1060.09          | 9433.2                      | -4.91(13)                                      |                                                    |                                                   |
| 17         | 1062.86          | 9408.6                      | -5.26(4)                                       | -6.00(39)                                          | -5.35(86)                                         |
| 18         | 1065.77          | 9382.9                      | -6.35(27)                                      |                                                    |                                                   |

Note: Asterisks indicate new peaks.

is helped by the fact that these transitions happen solely to the lowest energy state in  $^4I_{9/2}$  (e.g., Muñoz-Santiuste et al. 2017). Additionally, we are able to separate and identify the majority of the individual Stark levels due to differences in intensity, line shape and line width in both sets of transitions (Henderson et al. 1967). Once compressed above  $\sim 1$  GPa, a fifth band associated with the transition from  $^4H_{9/2} \rightarrow ^4I_{9/2}$  is observed (Fig. 7). As an aside, we note that the appearance of this sequence of luminescence bands is essentially indistinguishable in both amplitude and absolute position (laser frequency minus reported Raman frequency) from bands that have been attributed to Raman-active hydroxyl stretching vibrations observed under 633 nm excitation in bastnäsite from Pakistan (Frost and Dickfos 2007).

Around 2 GPa, the pressure-induced removal of accidental degeneracies in the higher energy transitions becomes apparent, and the appearance of new bands continues as pressure is increased in bastnäsite-I. For example, the intensity of peak 25 decreases, likely also from the loss of degeneracy. As with most other emission bands in this system, the nine bands associated with these transitions all shift to lower energy with pressure up to 25 GPa (Fig. 7).

Above 25 GPa, two new bands emerge, and the highest energy mode 23 undergoes an anomalous shift to higher energy with pressure from 25 to 38 GPa. This shift is consistent with the changes in the highest-lying bands associated with the  $^4F_{3/2} \rightarrow ^4I_{9/2}$  transitions. All other peaks shift to lower energy with pressure until 38 GPa; a representative deconvoluted spectrum at 31.2 GPa is shown in Supplemental<sup>1</sup> Figure S5c.

Bastnäsite-III has broader and lower intensity luminescence peaks for this set of transitions (Fig. 7). This drop in intensity and increase in breadth is consistent with there being an additional REE site within the unit cell of this phase, but the resolution of a separate full manifold of bands associated with the new site is not feasible. In addition to the intensity changes, there are discontinuous changes in some of the bands. For example, the lowest energy peak 30 becomes unresolvable above the transition, peak 25 shifts discontinuously to a higher energy, and a new band, (peak 26a\*\*), branches out from between peaks 26a and 26a\* (Fig. 7, Table 6); the deconvolution of the highest pressure spectrum is shown in Supplemental<sup>1</sup> Figure S6c. As with the other luminescent sets of transitions, the higher pressure phase change involves more marked changes in the

spectra than the lower pressure transition from bastnäsite-I to bastnäsite-II.

### R1-R2 separation

The energy differences between the R1 and R2 bands are perhaps most readily associated with structural changes in the REE site under pressure (Fig. 8). While the absolute shift of each energy level with pressure is quite small (Fig. 5), the separation in energy between these two bands does change. An increase in energy separation is characteristic of a REE site that is decreasing in symmetry/increasing in distortion, and a decrease indicates the converse: a transition toward higher symmetry/lower distortion (e.g., Hua et al. 1996). From 0 to 15 GPa, the difference between these two modes initially decreases indicating a REE site that is becoming more symmetric or, in other words, R1 and R2 are becoming closer to degenerate (Fig. 8). At 15 GPa, the difference between these two energy levels increases with pressure until 25 GPa, thus indicating that the REE site is becoming more distorted. Above 25 GPa, the change in energy level separation is negligible until 38 GPa, where the onset of a new REE site occurs. The energy level splitting of the new and old REE site differ dramatically. The most striking change is the energy difference between R1 and R2: at 50 GPa, the difference between R1 and R2 in the original REE site is  $\sim 42$  cm<sup>-1</sup>, and in the new site, the difference is  $\sim 96$  cm<sup>-1</sup>. While this difference is large, an energy difference between R1 and R2 of above 100 cm<sup>-1</sup> is common within ambient pressure, highly distorted Nd<sup>3+</sup> sites. For example, the R2-R1 separation

**TABLE 6.**  $^4F_{5/2} + ^2H_{9/2} \rightarrow ^4I_{9/2}$  luminescence as a function of pressure in bastnäsite

| Assignment           | $\lambda_0$ (nm) | $\nu_0$ (cm <sup>-1</sup> ) | Bastnäsite-I<br>to 25 GPa<br>(cm <sup>-1</sup> /GPa) | Bastnäsite-II<br>25 to 40 GPa<br>(cm <sup>-1</sup> /GPa) | Bastnäsite-III<br>4 to 50 GPa<br>(cm <sup>-1</sup> /GPa) |
|----------------------|------------------|-----------------------------|------------------------------------------------------|----------------------------------------------------------|----------------------------------------------------------|
| 23 ( $^2H_{9/2}$ )   | 791.46           | 12634.9                     | -0.37(8)                                             | 0.57(9)                                                  | -0.49(18)                                                |
| 24 ( $^2H_{9/2}$ )   |                  |                             | -0.04(7)                                             | -0.73(7)                                                 | -1.30(13)                                                |
| 24* ( $^2H_{9/2}$ )  |                  |                             | -1.52(33)                                            | -2.01(12)                                                | -1.90(21)                                                |
| 25* ( $^2H_{9/2}$ )  |                  |                             |                                                      |                                                          | -2.00(24)                                                |
| 25 ( $^2H_{9/2}$ )   | 793.65           | 12600.0                     | -1.94(6)                                             | -2.51(11)                                                |                                                          |
| 26a ( $^2H_{9/2}$ )  |                  |                             | -2.30(25)                                            | -0.80(19)                                                | -0.60(23)                                                |
| 26a* ( $^2H_{9/2}$ ) |                  |                             |                                                      | -1.83(15)                                                | -2.04(33)                                                |
| 26a**                |                  |                             |                                                      |                                                          | -2.00(37)                                                |
| 26* ( $^2H_{9/2}$ )  |                  |                             |                                                      | -3.72(29)                                                | -0.73(70)                                                |
| 26 ( $^2H_{9/2}$ )   | 798.04           | 12530.7                     | -5.05(47)                                            | -3.20(11)                                                | -3.08(75)                                                |
| 27 ( $^2H_{9/2}$ )   | 803.34           | 12448.0                     | -3.82(18)                                            | -3.19(37)                                                | 2.79(28)                                                 |
| 28 ( $^4F_{5/2}$ )   | 807.94           | 12377.2                     | -4.41(68)                                            | -3.93(54)                                                | -3.16(76)                                                |
| 29 ( $^4F_{5/2}$ )   | 814.90           | 12271.4                     | -5.98(32)                                            | -4.36(21)                                                | -0.93(43)                                                |
| 30 ( $^4F_{5/2}$ )   | 822.71           | 12155.0                     | -7.22(14)                                            | -4.99(22)                                                |                                                          |

Note: Asterisks indicate new peaks.

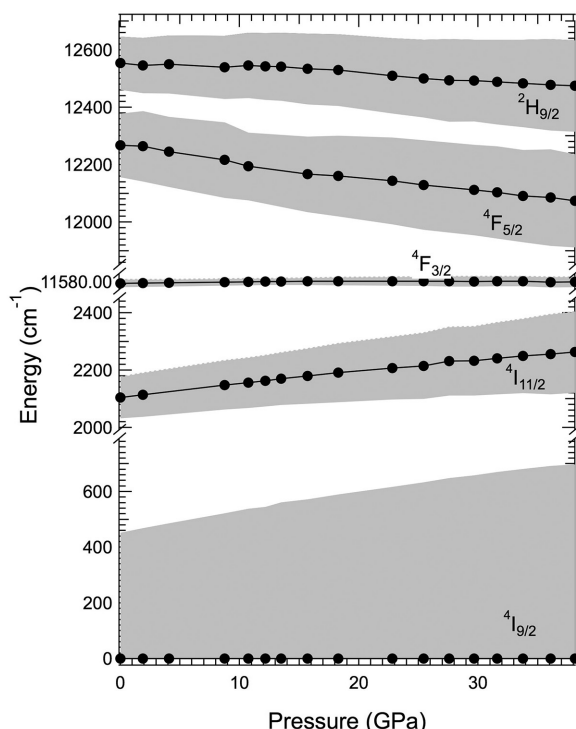


**FIGURE 9.** Energy levels of Nd<sup>3+</sup> with pressure; shaded regions designate the observed spread in energy levels, and lines represent the barycenters of the Stark levels, with the exception of the  $^4I_{9/2}$  level, where the line represents the ground state of the system.

is 123 cm<sup>-1</sup> in CAZGAR (CaZn<sub>2</sub>Y<sub>2</sub>Ge<sub>3</sub>O<sub>12</sub>; Sardar and Yow 1998) and 116 cm<sup>-1</sup> in CaY<sub>2</sub>MgGe<sub>3</sub>O<sub>12</sub> (Sharp et al. 1974). In addition, if site distortion increases under pressure, increased separation is expected, so this doubling of the R2-R1 separation in the new site to nearly 100 cm<sup>-1</sup> is not outside the bounds of sites observed at ambient pressure in other phases.

### Separation of energy levels with pressure

Over most of the pressure range (15–50 GPa) of this experiment, the luminescence bands within each set of transitions separate under compression, indicating a progressive increase in the crystal field splitting, which is likely associated with a pressure-induced increase in the distortion of the REE site. The change in the separation energy of <sup>4</sup>I<sub>9/2</sub> (Z1–Z5) is 6.03 cm<sup>-1</sup>/GPa, <sup>4</sup>I<sub>11/2</sub> (Y1–Y4), is 3.61 cm<sup>-1</sup>/GPa, <sup>4</sup>F<sub>3/2</sub> (R1–R2) is 0.32 cm<sup>-1</sup>/GPa, <sup>4</sup>F<sub>5/2</sub> is 2.81 cm<sup>-1</sup>/GPa, and <sup>2</sup>H<sub>9/2</sub> is 3.63 cm<sup>-1</sup>/GPa. Figure 9 displays the overall splitting of the spin-orbit coupling energy levels, with the barycenters (the arithmetic means) shown as the energy levels (for all but the ground state). Similar rates of changes in the separation of the energy levels <sup>4</sup>I<sub>9/2</sub>, <sup>4</sup>I<sub>11/2</sub>, and <sup>4</sup>F<sub>3/2</sub> have been observed previously in YAlO<sub>3</sub>:Nd at high pressures (Hernández-Rodríguez et al. 2018a). Several distinct and separable effects are present in Figure 9. First, the net separation between the Stark levels of the <sup>2</sup>H, <sup>4</sup>F, and <sup>4</sup>I levels decreases under compression. As shown in Figure 3, this indicates that the splitting of the f-levels induced by Coulombic splitting is reduced under compression. This implies that the interaction between f-orbitals is lessened, indicating that covalency of the REE site in bastnäsite increases under compression (this is a manifestation of an increased nephelauxetic effect). The shift in Coulombic splitting can be semi-quantitatively assessed from Figure 9: the separations of the averages of the H, F, and I levels decrease by ~2% for both the H to I level differences and the F to I level differences between 0 and 40 GPa. The relative tradeoffs between a pressure-induced contraction in the size of the f-orbitals relative to decreased electron localization in these levels induced by increased covalency in generating this change in splitting by Coulombic repulsion are unclear. If, as an end-member upper bound on the maximum shift in electron occupancy, the entire shift in Coulombic splitting is attributed to a change in occupancy, then a decrease of ~1% in electron occupancy in the f-orbitals is indicated by the results of Figure 9: a relatively small, but spectroscopically significant, shift relative to the magnitude of the change in volume of the system (which is of order ~19% at 40 GPa using Rowland's (2017) bulk modulus value). In passing, we note that this semi-quantitative estimate, derived from luminescence spectra, could be rigorously tested via single-crystal diffraction constraints on the pressure-dependence of electron density distributions.

In terms of the spin-orbit splittings (Fig. 3), the separation between the spin-orbit split levels decreases for the <sup>4</sup>F levels but may increase modestly for the <sup>4</sup>I levels (Fig. 9). A decreased electron density induced by enhanced covalency would be generally anticipated to decrease spin-orbit coupling, as is clearly the case for the <sup>4</sup>F level. We have no simple explanation for the enhanced spin-orbit coupling of the <sup>4</sup>I level. Last, the splittings of the Stark levels are expected to be enhanced by both increased distortion and enhanced crystal field strength (e.g., Tröster 2003). This is

clearly observed over the pressure range of our experiments, with the spread between levels (the shaded regions in Fig. 9) clearly expanding under pressure. Indeed, the maximum value of Stark splitting has been shown to scale with the crystal field strength at the site (Auzel and Malta 1983).

### Anomalous pressure shifts in Nd<sup>3+</sup> luminescence

Overall, the general expectation is that luminescence lines will shift to lower energy with pressure; as pressure is increased, the ligands bound to the Nd<sup>3+</sup> ion are brought closer to the Nd<sup>3+</sup> ion, and all electronic levels move to higher energy, but the overarching trend is that higher lying levels tend to shift less quickly to higher energies than lower levels. Thus, the effect of the change in spatial distance between the REE and ligands causes the energy difference between the ground state and the excited states to become lessened with compaction. Previously, anomalous behavior has been observed in the perovskite structured Nd<sup>3+</sup>-doped YAlO<sub>3</sub>: blueshifts of a subset of peaks associated with the transition from <sup>4</sup>F<sub>3/2</sub> to <sup>4</sup>I<sub>9/2</sub> have been observed (Barnett et al. 1973; Hua and Vohra 1997; Hernández-Rodríguez et al. 2018a). These anomalous shifts have been attributed to a combination of increasing crystal field strength (and hence splitting) and hybridization of the f-orbitals (Hernández-Rodríguez et al. 2018a). The vast majority of the luminescence bands shift to lower energy with pressure in bastnäsite. However, two luminescence bands have pressure shifts that involve an increase in energy in bastnäsite-II (<sup>2</sup>H<sub>9/2</sub> → <sup>4</sup>I<sub>9/2</sub>, peak 23; and <sup>4</sup>F<sub>3/2</sub> → <sup>4</sup>I<sub>9/2</sub>, peak 3), and two increase in bastnäsite-III: the new mode associated with R2 peak 1 and peak 6 in <sup>4</sup>F<sub>3/2</sub> → <sup>4</sup>I<sub>9/2</sub>. Notably, these bands are among the higher energy lines associated with each of their respective transitions; hence, the role of enhanced Stark splitting counterbalances the general decrease in energy of the barycenter of their respective transitions. Thus, these atypical shifts can be attributed to complex crystal field interactions between the Nd<sup>3+</sup> ion and O/F ligands induced by an enhanced site distortion, and hence anisotropic compression within the REE site.

### IMPLICATIONS

The bonding of both the carbonate units and neodymium ions under pressure is constrained in this critically important ore mineral. Notably, bastnäsite undergoes two apparently distortional phase transitions under compression: the first, near 25 GPa, appears to involve primarily a shift in the symmetry of the carbonate ion, while the second, which initiates at 38 GPa, generates a second REE site in this material. From an overarching perspective, the overall layered bonding of the bastnäsite structure, with REE-F layers and carbonate layers, appears remarkably stable under compression: no indication is seen within this pressure range of any weakening of C–O bonds that would indicate an incipient increase in coordination, or onset of dimerization associated with the carbonate group (e.g., Vennari et al. 2018). The general structural stability of this phase indicates that REE retention in the Earth's mantle may continue to involve bonding within phases containing both halogens and carbonate ions. Hence, REE element cycling, which has been used to constrain mantle differentiation, may critically depend on the presence of trace carbonate- and/or halogen-bearing phases at depth. Thus, the affinity of REE for carbonate-associated lithologies in the near surface may persist to depth within the planet.

From a high-pressure science perspective, bastnäsite (which is quite stable under compression at 300 K) could represent a candidate for a high-pressure luminescent calibrant in the near IR region from 0 to ~38 GPa: the transition from R1 to Y2 (Band 14), which is an intense mode in the NIR region does not split under pressure and shifts robustly and continuously through the transition from bastnäsite-I and bastnäsite-II. Furthermore, the splitting between two of the other intense peaks (R1 and R2) may provide a sensitive probe of the degree of distortion of REE sites within a range of REE-bearing oxides and halides under compression.

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## Endnote:

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